Towards a new VLBI system for geodesy and astrometry

J. Böhm⁽¹⁾, J. Wresnik⁽¹⁾, H. Schuh⁽¹⁾, D. Behrend⁽²⁾

⁽¹⁾ Vienna University of Technology⁽²⁾ NVI/NASA Goddard Space Flight Center





TECHNISCHE UNIVERSITÄT WIEN

VIENNA UNIVERSITY OF TECHNOLOGY



Very Long Baseline Interferometry

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- CRF
- Celestial Pole
- UT1 UTC
- Primary technique for
 - EOP (complete set of parameters)
 - TRF (most precise for long baselines, scale)

IVS - International VLBI Service for Geodesy and Astrometry

IVS is a Service of

- IAG (International Association of Geodesy)
- IAU (International Astronomical Union)
- FAGS (Federation of Astronomical and Geophysical Data Analysis Services)

IVS Goals

- provide a service to support geodetic, geophysical and astrometric research and operational activities
- promote research and development in the VLBI system
- interact with the community of users of VLBI products and integrate VLBI into a global Earth observing system

IVS Component Map



Radio telescopes: Wettzell (20m), Effelsberg (100m)



IVS Products (1) ICRF (Ext. 1, Ext. 2)



Figure 1: Distribution of the 608 ICRF sources on an Aitoff equal-area projection of the celestial sphere. The dotted line represents the Galactic equator.

IVS Products (2) dUT1 = UT1 - UTC



IVS Products (3) Scale for the ITRF2005

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Altamimi et al. (2007)

IVS Products (4) Baseline lengths



IVS VLBI2010 WG 3 Goals

- 1 mm position accuracy
- 0.1 mm/year velocity accuracy
- continuous
 measurement of EOP
- rapid generation and distribution of IVS products



IVS VLBI2010 WG3 Recommendations

- Design a new observing system based on small antennas (12m), fast moving, operated unattended, mechanically reliable
- Broad continuous frequency range (2-15 GHz) which includes S- and X-Band
- Upgrade of large antennas to preserve continuity
- Transfer data with combination of high speed networks and high rate disk systems
- Examine the possibility for a new correlator system (software correlator?)

IVS VLBI2010 Committee

- Established September 2005
- Promote and guide research into the improvement of the "technique" of geodetic VLBI
- Take an integrated view of VLBI, evaluate effectiveness of proposed system changes with respect to final products
- Take responsibility for encouraging the implementation of the recommendations of WG3

VLBI2010 Simulations



VLBI2010 Simulations



Monte Carlo Simulations

 $o - c = (wzd_2 \cdot mfw_2(e) + cl_2) + (wzd_1 \cdot mfw_1(e) + cl_1) + wn_{Bsl}$

wet zenith delays: random walk PSD: 0.1 psec²/s 0.7 psec²/s turbulence model (Onsala, Sweden) **clocks:** ASD 1·10⁻¹⁴@50min 2·10⁻¹⁵@15min

observation error: 4, 8, 16 psec

Simulate 25 identical 24 hour sessions

Wet delay is the limiting factor



Slew rate studies

_ ا	×	azim.	elev.
	16 stations		
Repeatability [mm]	2·10 ⁻¹⁵ @15min clocks		
	4 psec white noise	1.5°/s	0.7°/s
	turbulence model	3.0°/s	0.7°/s
		4.5°/s	2.1°/s
		6.0°/s	2.1°/s
		12.0°/s	3.5°/s
0			
0	baseline length [1000 km] 13	3	

VLBI2010 is supporting several proposals (1)

- Korean Institutes (KASI, NGII, Ajou University)
- Geoscience Australia: proposal for 3 fundamental stations
- Univ. Tasmania (Hobart): getting operation money
- University of Concepción (Chile): developing telescope
- ISRO, India: 32m telescope for lunar mission extended for geodetic VLBI
- NASA Haystack: support of VLBI 2010 telescope

VLBI2010 is supporting several proposals (2)

BKG: twin-telescope 2008-2010



- higher observation density
 - continuous observations
 - better determination of systematic effects (one clock)
- one more local tie

observation station

Summary -Role of IVS in GGOS

Contributes unique strength of VLBI technique

- CRF products (quasar positions)
- Complete set of EOP parameter; uniquely UT1-UTC and nutation; provides link CRF ↔ TRF
- TRF products (station positions, velocities, scale, baseline lengths)
- Tropospheric and ionospheric parameters
- other geodynamical and astronomical parameters (e.g., Love numbers, relativistic parameters)
- Support for satellite tracking

Thanks for your attention!



IVS Products (5) Station coordinates / velocities



(from a global TRF solution derived by the DGFI, Munich)